#### **Cosmic acceleration from structure formation**

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### Outline

#### 1. Statement:

volume-averaged expansion of space can accelerate even if the local expansion decelerates everywhere

- 2. Mathematical foundation
- 3. Physical interpretation
- 4. Observational status
- 5. Conclusions

# This talk is based on General Relativity. No new physics is proposed.



• Raychaudhuri equation in an irrotational dust universe:

$$\frac{\partial}{\partial t}\theta = \underbrace{-\sigma^2 - \frac{1}{3}\theta^2 - \rho}_{\leq 0}$$

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• For large variance of the expansion rate, the right hand side can be positive

## Numerical proof by Chuang, Gu and Hwang using an exact LTB solution



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Short answer: the volume becomes dominated by the fast-expanding regions



### Volume $V_0$ Volume $V_0$













 $\implies$  expansion has accelerated from 0 to h

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  - $-\langle (\theta \langle \theta \rangle)^2 \rangle$  not enough to accelerate volume-averaged expansion, but...
  - relating the volume-averages to local observables leads to apparent acceleration

• Averages of inhomogeneous quantities do not evolve in time like the corresponding homogeneous quantities

• Average expansion can accelerate even if local expansion decelerates everywhere

• Structure formation may account for the observed cosmic acceleration within standard physics

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